



Presented at the COMSOL Conference 2009 Boston

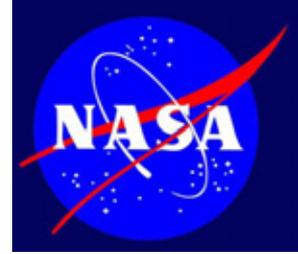
# Using Microwaves for Extracting Water from the Moon

**COMSOL Conference**  
**October 9, 2009**  
**Boston, MA**

**Edwin Ethridge, PhD, Principal Investigator**  
*Materials & Processes Laboratory*  
*NASA-MSFC-EM40, Huntsville, AL*

**William Kaukler, PhD**  
*Department of Chemistry*  
*The University of Alabama in Huntsville*

**Frank Hepburn**  
*Materials & Processes Laboratory*  
*NASA-MSFC-EM20, Huntsville, AL*



# **Water is one of the Most Plentiful Compounds in the Universe**

- **Earth's water came from comets early in its history.**
- **Mars has vast quantities of water not only at the poles but also at lower latitudes ( $>55^{\circ}$ ).**
- **Significant quantities of water are present on several moons of Jupiter (Europa, Ganymede, and Callisto), Saturn (Enceladus), and Neptune (Triton).**

# Chronology of “Water on the Moon”



**1959**, prior to Apollo, scientists **speculated** that there should be some residual **water on the moon from cometary impacts**.

**Apollo** found “**no water**”.

**1994**, the SDI-NASA **Clementine spacecraft** mapped the surface. **Microwave radio signals from South pole shadowed craters were consistent with the presence of water**.

**1998 – Prospector - Neutron Spectrometer – high H concentrations at poles**

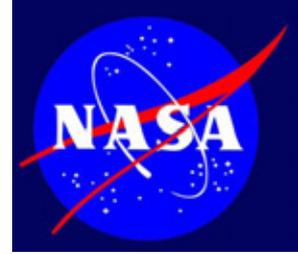
**September 2009 – water observed at diverse locations of moon**

**Chandrayaan-1 - Moon Mineralogy Mapper (M3)**

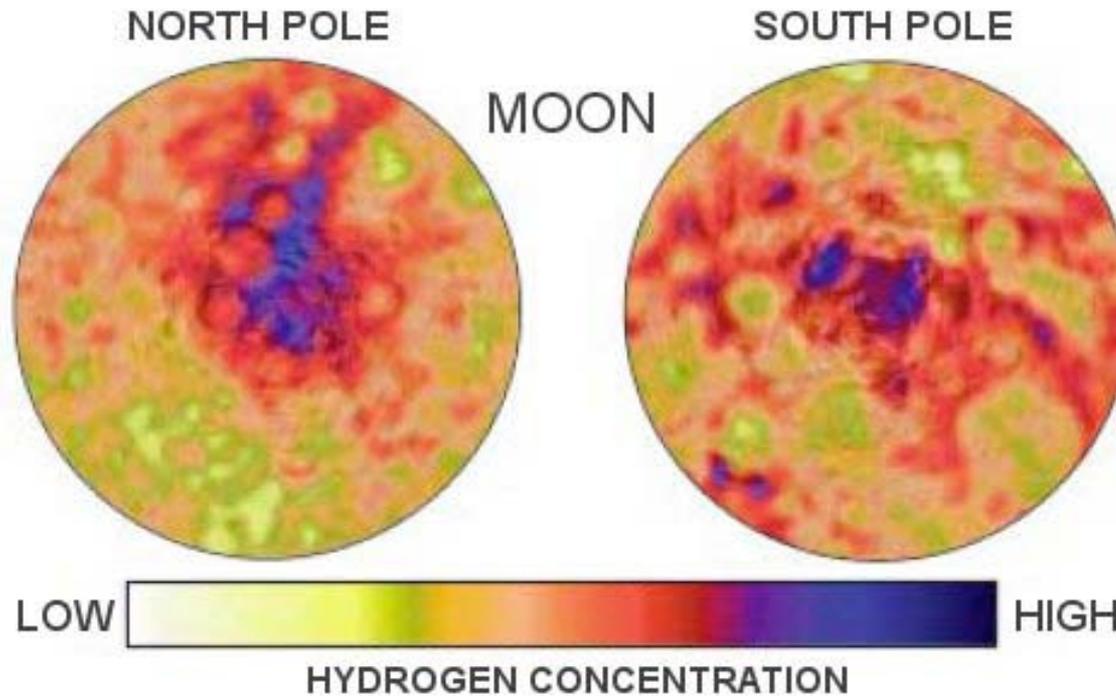
**Cassini - Visual and Infrared Mapping Spectrometer (VIMS)**

**EPOXI spacecraft - High-Resolution Infrared Imaging Spectrometer**

**October 9, 2009 - Lunar CRater Observation and Sensing Satellite - LCROSS**



# 1998- Lunar Prospector Hydrogen Maps of the Lunar Poles



**1998 Lunar Prospector - Neutron Spectrometer** scanned for hydrogen-rich minerals. **Polar craters** yielded **neutron ratios** which **indicated hydrogen => H<sub>2</sub>O**

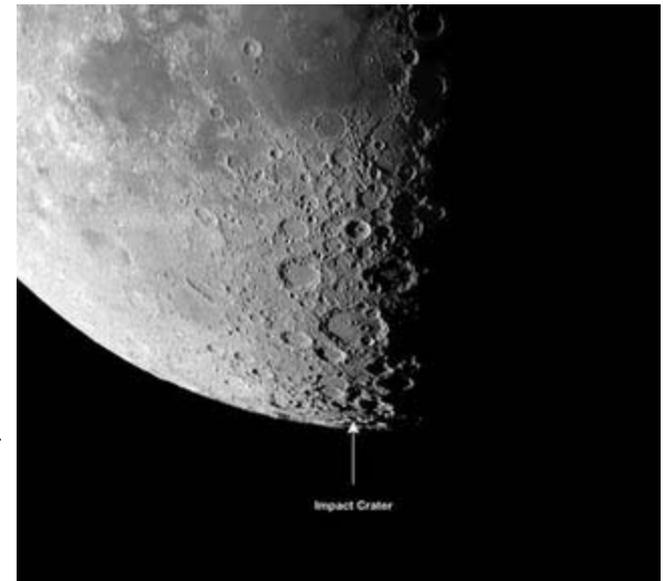
**Average H<sub>2</sub>O concentration ~2% → Theoretically approaching Billions of tons**



# Lunar CRater Observation and Sensing Satellite - LCROSS

- **THIS MORNING** the **Centaur upper stage impacted** a permanently shadowed crater, Cabeus A near the south pole of the Moon.
- **Mission Objective - confirm the presence of water ice** in a permanently shadowed crater at the **Moon's South Pole**.
- **Spectral analysis** of the resulting impact plume will look for **water ice**.

Crater Cabeus A →

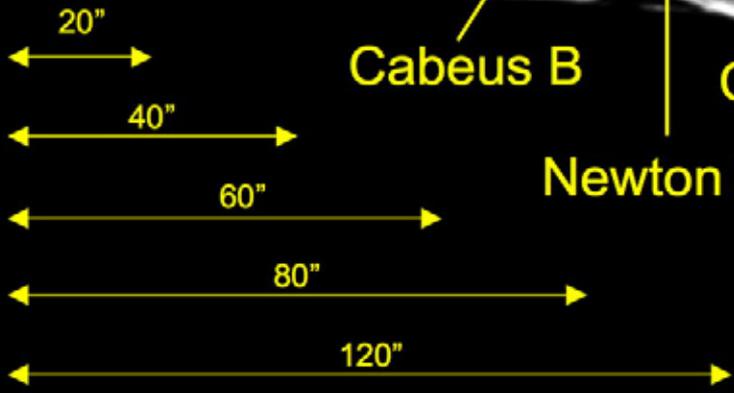
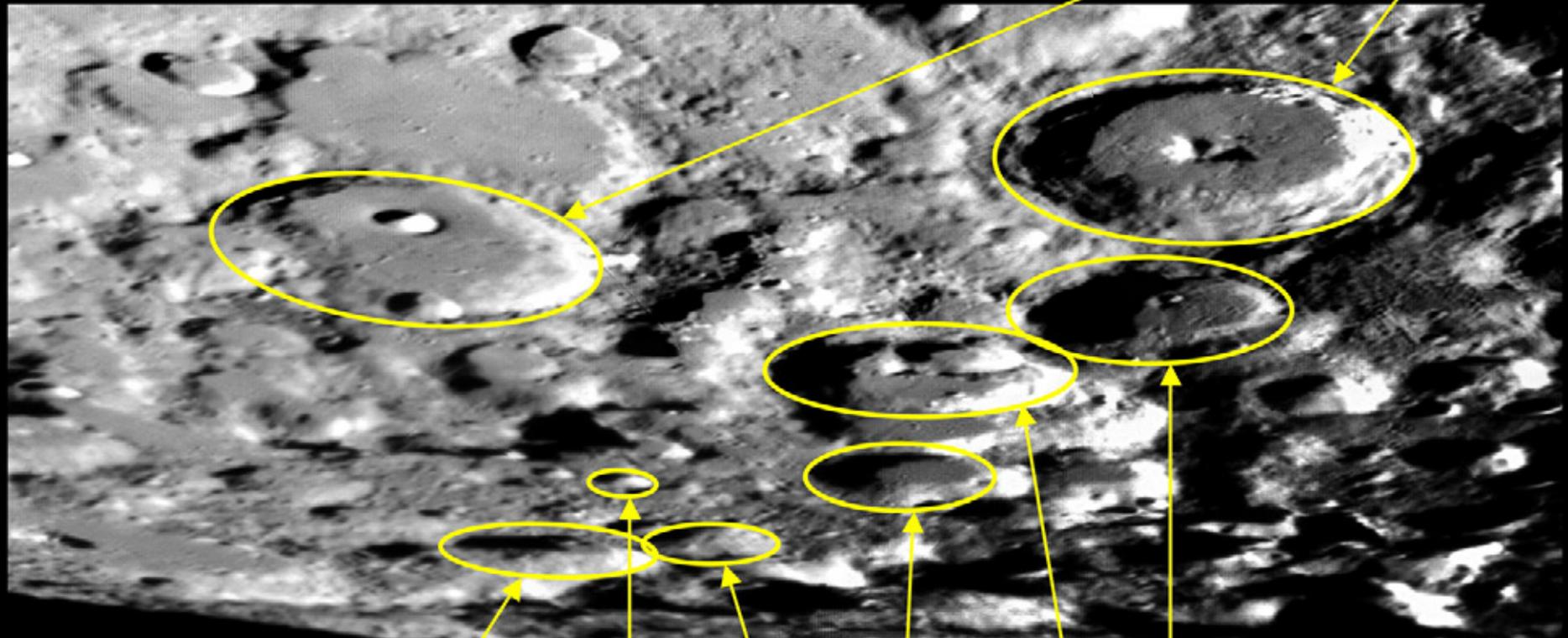


# Lunar South Pole Map

Medium field

Casatus

Moretus



Cabeus B

Cabeus A

Newton E

Newton A

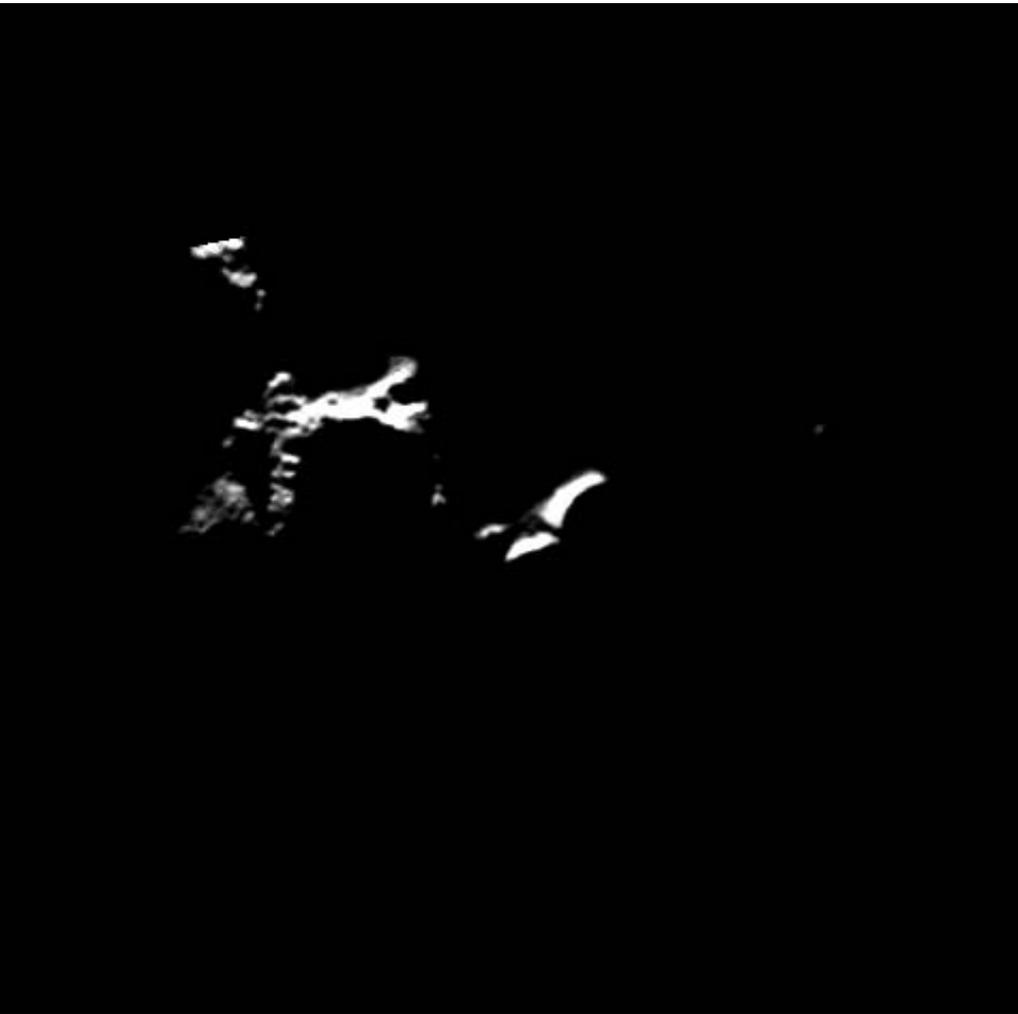
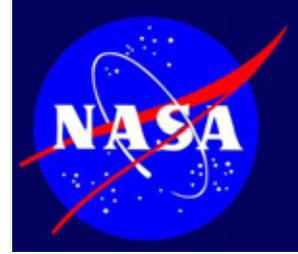
Newton

Short

NMSU / MSFC  
Tortugas Observatory 24"  
0.9 - 1.7  $\mu\text{m}$  InGaAs Camera

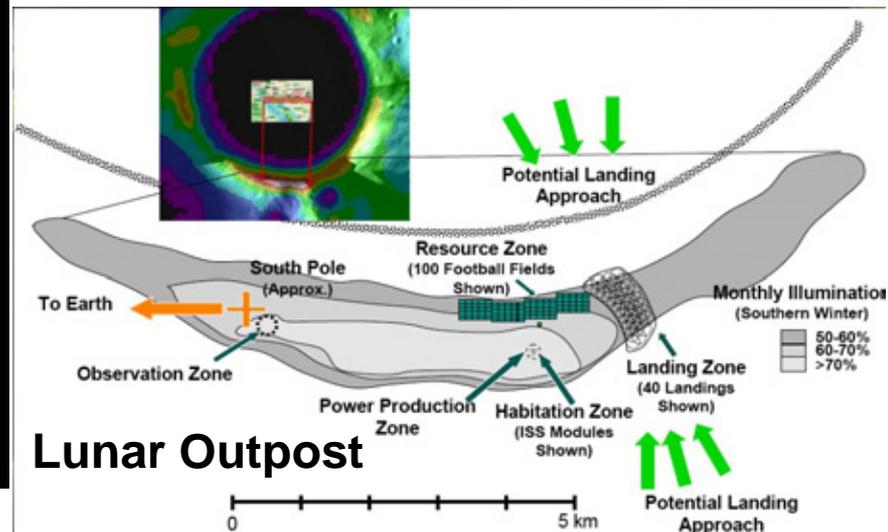
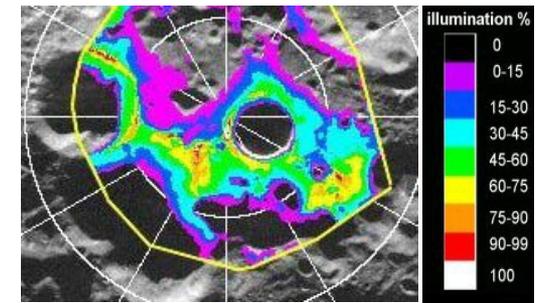
Scale at time of LCROSS impact (1.78 km/arcsec)

# Illumination of the South Polar Region of the Moon over one Lunar Day (28 days)



<http://www.lpi.usra.edu/resources/clemen/clemen.html>

## Moon Polar light map



**Continuous Sunlight, Continuous Contact with Earth, “Moderate” Temperature, and Resources (water ice?)**

# Importance of Water

- **Water (and oxygen) for the manned lunar outpost.**  
**Resupply requirements:**  
**1 ton water and 1 ton oxygen per year**  
**~\$10,000 to \$100,000/ pound from Earth to moon**  
**Increasing to 10 tons per year**
- **Eventually (perhaps) - rocket propellant for exploration,**  
**Electrolyze water into H and O,**  
**liquify to LOX and LH**  
**10's of tons not launched out of Earth's**  
**gravitational well.**

# Water on the Moon

**The water on the moon is in a number of different forms.**

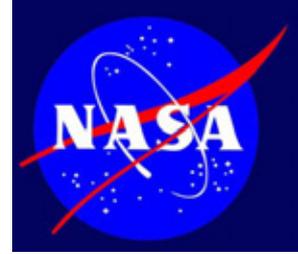
**Chemically Bound water – Hydroxyl groups -OH**

- **The hydroxyls can be on the surface of the grains of soil.**
- **They can also be part of the chemical composition of the rock.**
- **This could include other hydrated chemical species (at the poles).**
- **Relatively small weight percent measured in Apollo soils.**

**Molecular water H<sub>2</sub>O**

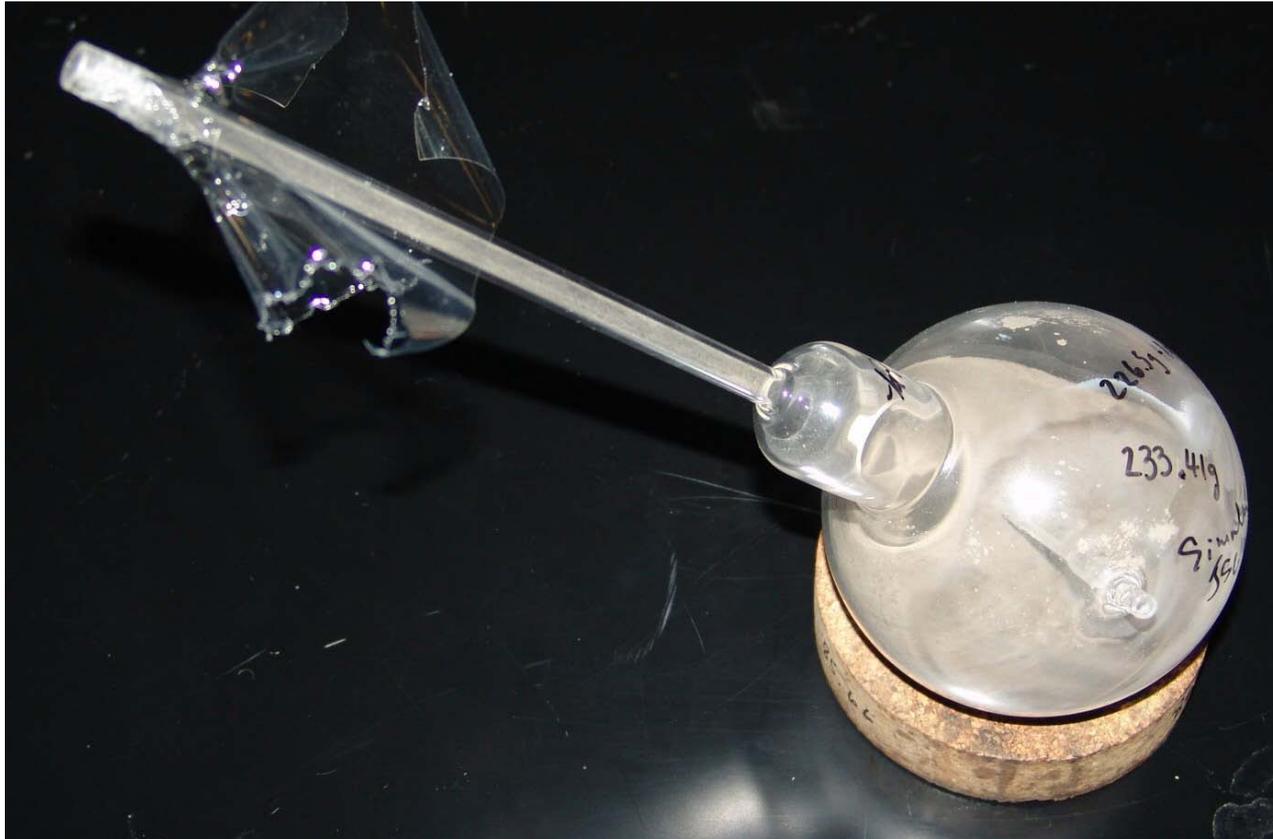
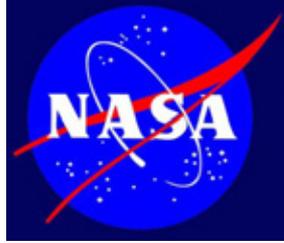
- **Recently measured water vapor near the lunar surface.**
- **Water molecules lightly bound to Si<sup>+</sup> dipoles on the surface of grains of lunar soil.**
- **Water ice physically condensed at poles. This physically condensed cryogenically trapped water ice is speculated to be present in relatively high concentrations (on the order of 2 weight percent),**  
**This is the water we have proposed to extract with microwaves.**

# Use of Microwaves for the Extraction of Lunar Water

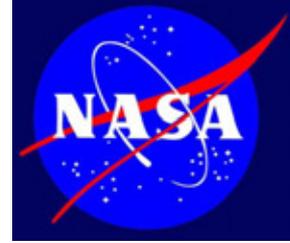


- **Lunar Soil (in vacuum) is a Super Thermal Insulator  
Is like an aerogel, very low heat flow.**
- **Microwave energy penetrates the soil heating from the  
inside out. Penetration depth is dependent  
on Frequency and dielectric properties.**
- **Conversion from electricity through microwaves to heat,  
efficient, heating causes sublimation of water ice.**
- **Excavation may not be required,  
Cryogenic water ice is as hard as granite  
Saving energy, infrastructure, and equipment  
Little if any disruption of lunar dust (hazard)**

# “Moon in a Bottle” Laboratory Proof of Principle

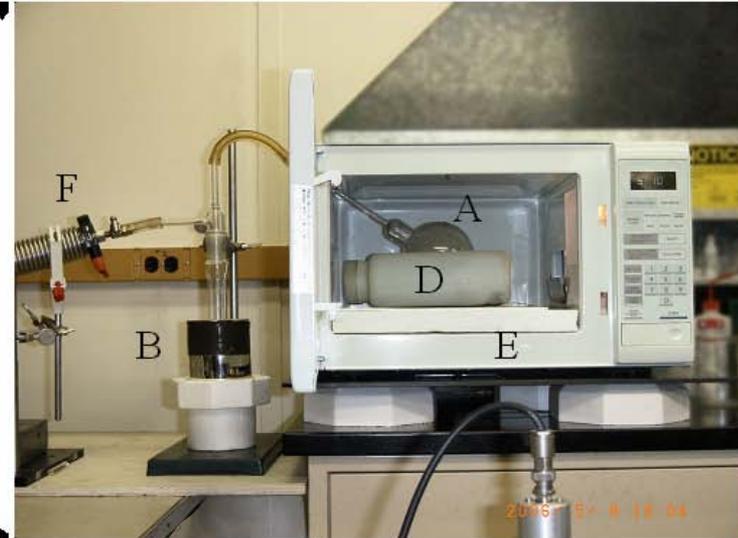
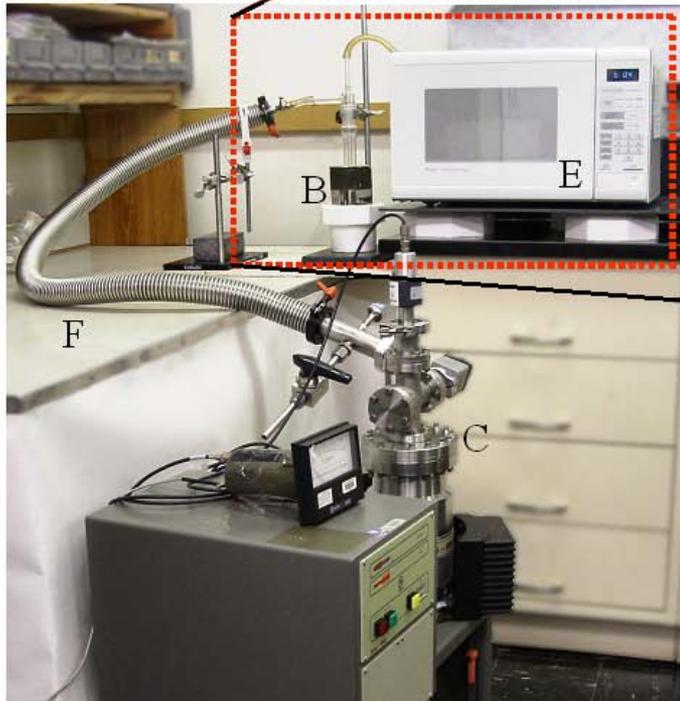


**Fused silica vessel with lunar permafrost simulant.**



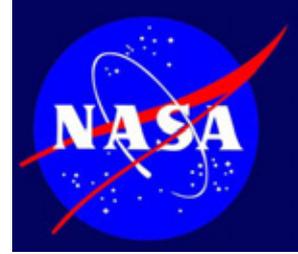
# Experimental Facilities

- Standardized lunar regolith simulant (JSC-1A), particle size distribution and chemistry of (Apollo 14).
- Water ice concentration (2 weight %, 2g)
- Temperature (-196 to -50C), LN2,
- Vacuum level (10<sup>-5</sup> torr),



- Bench top microwave facility**
- A. Vacuum quartz lunar regolith simulant vessel
  - B. Liquid nitrogen cold-trap
  - C. Turbo-molecular vacuum pump
  - D. LN2-cooled regolith simulant
  - E. Microwave oven chamber

# Extraction Efficiency

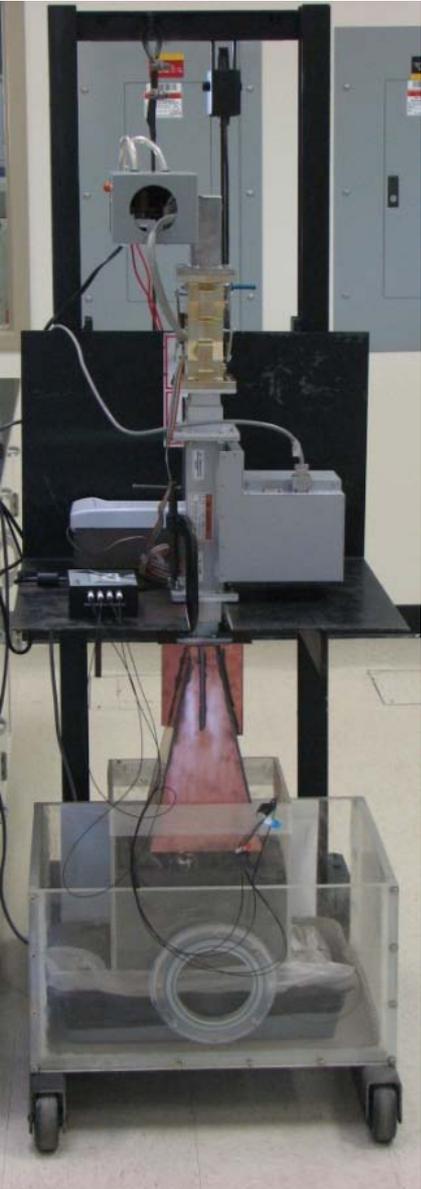


- **Microwaves coupled well to soil simulant at LN2 temperature.**
- **The regolith and the cold trap were weighed before and after the experiment.**
- **At least 95% of the water added to the regolith simulant was extracted (in 2 minutes) all below 0°C.**
- **Of the extracted water 99% was captured in the remote cold trap.**

# Microwave Lunar Water Extraction Prototype

- Magnetron source (2.45 GHz, 1100 W) with isolator, auto-tuner and copper high-gain horn.
- Mounting provides mobility over surface and height adjustment of horn.
- Temperatures within the bed of simulant (JSC-1A) were made using fiber optic temperature sensor in place during heating.

- Vacuum chamber evaluation of the microwave penetration and water vapor permeability through lunar soil simulant. -->



# Attenuation - Beer-Lambert law

## Penetration Depth\*

$$\text{Penetration} \sim \frac{\lambda (\epsilon')^{1/2}}{2 \pi \epsilon''}$$

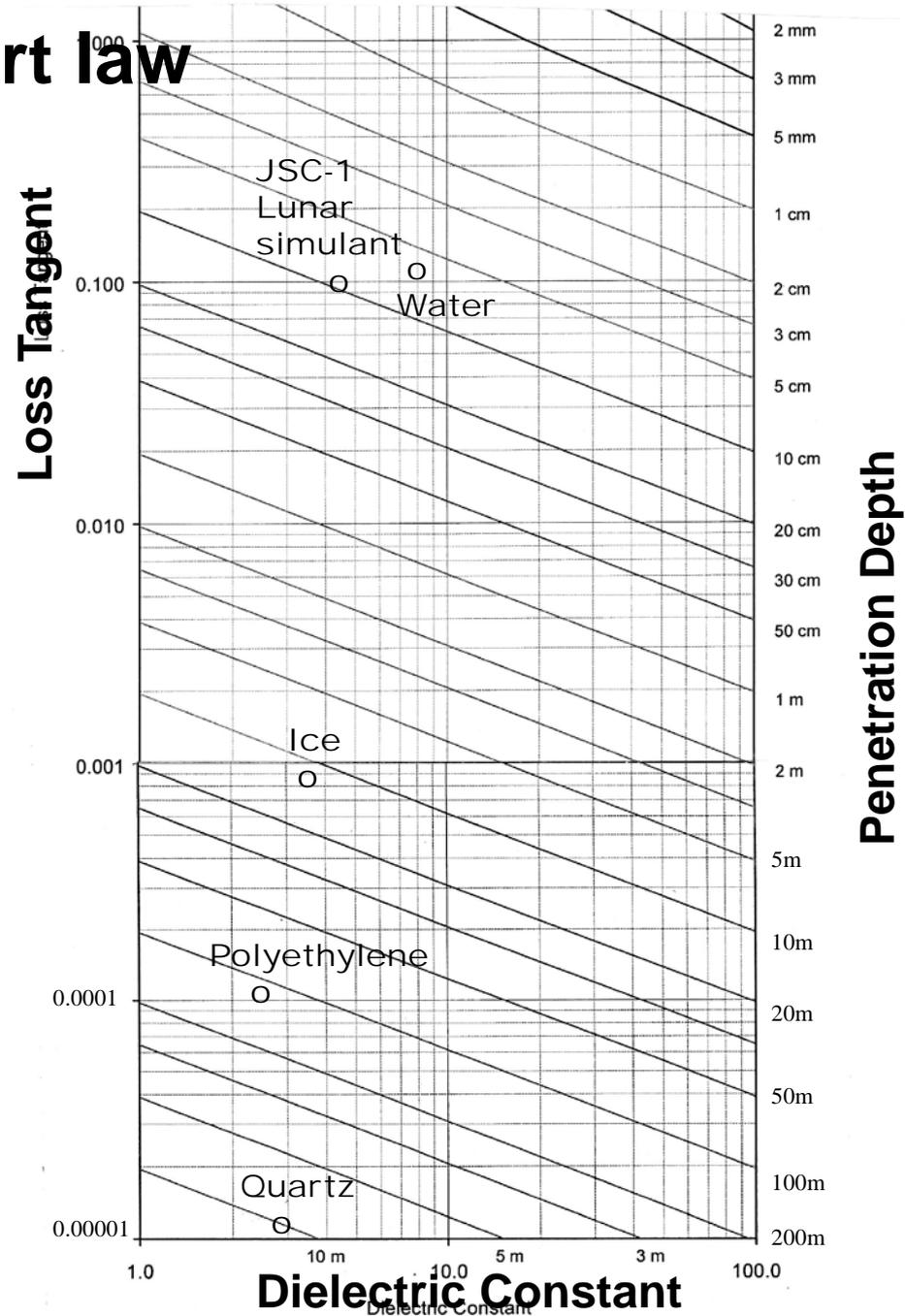
	$\epsilon''$	$\epsilon'$	Penetration
Water	0.1	8	6cm
Simulant JSC-1	0.1	4	10cm
Ice	0.001	3	1m
Polyethylene	0.0001	2	100m
quartz	0.00001	3	1000m

$$\epsilon = \epsilon' + j \epsilon''$$

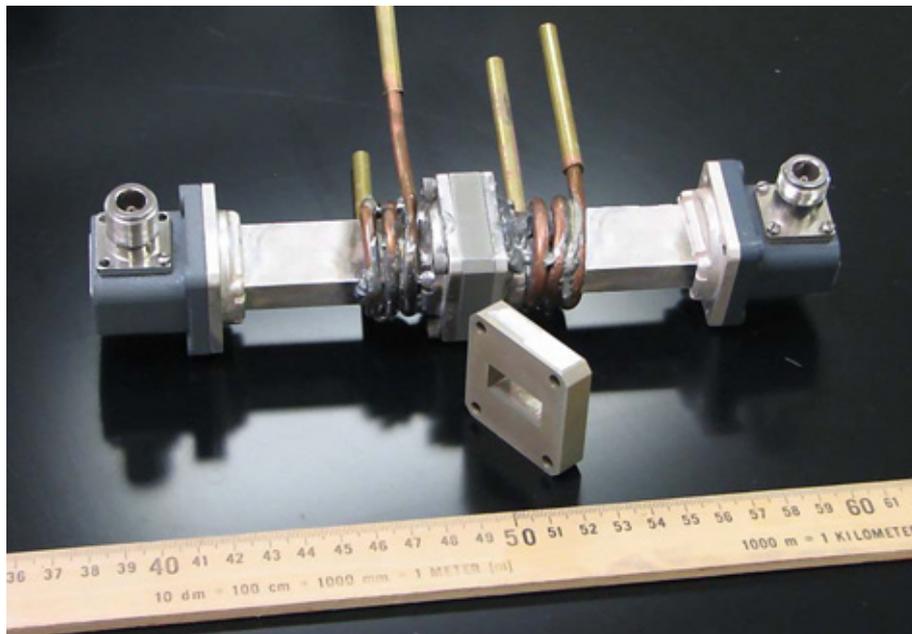
$\epsilon'$  dielectric constant

$\epsilon''$  loss factor

\*1/e (0.37) of the original value.



# Dielectric Property Measurements Lunar Soil Simulant – JSC-1A



With Frank Hepburn - EM20

- Our custom fabricated 10 GHz (range 8 – 12 GHz) X-band waveguide apparatus for dielectric measurements over a range of temperatures, LN<sub>2</sub> to above room temperature.
- Heating coils near the coax connectors (not shown) keep the instrument connections at room temperature while the sample residing between the cooling the coil is chilled with free-flowing LN<sub>2</sub>.

# Dielectric Properties of JSC-1A Simulant X-band 10 GHz, room temperature



Real and imaginary

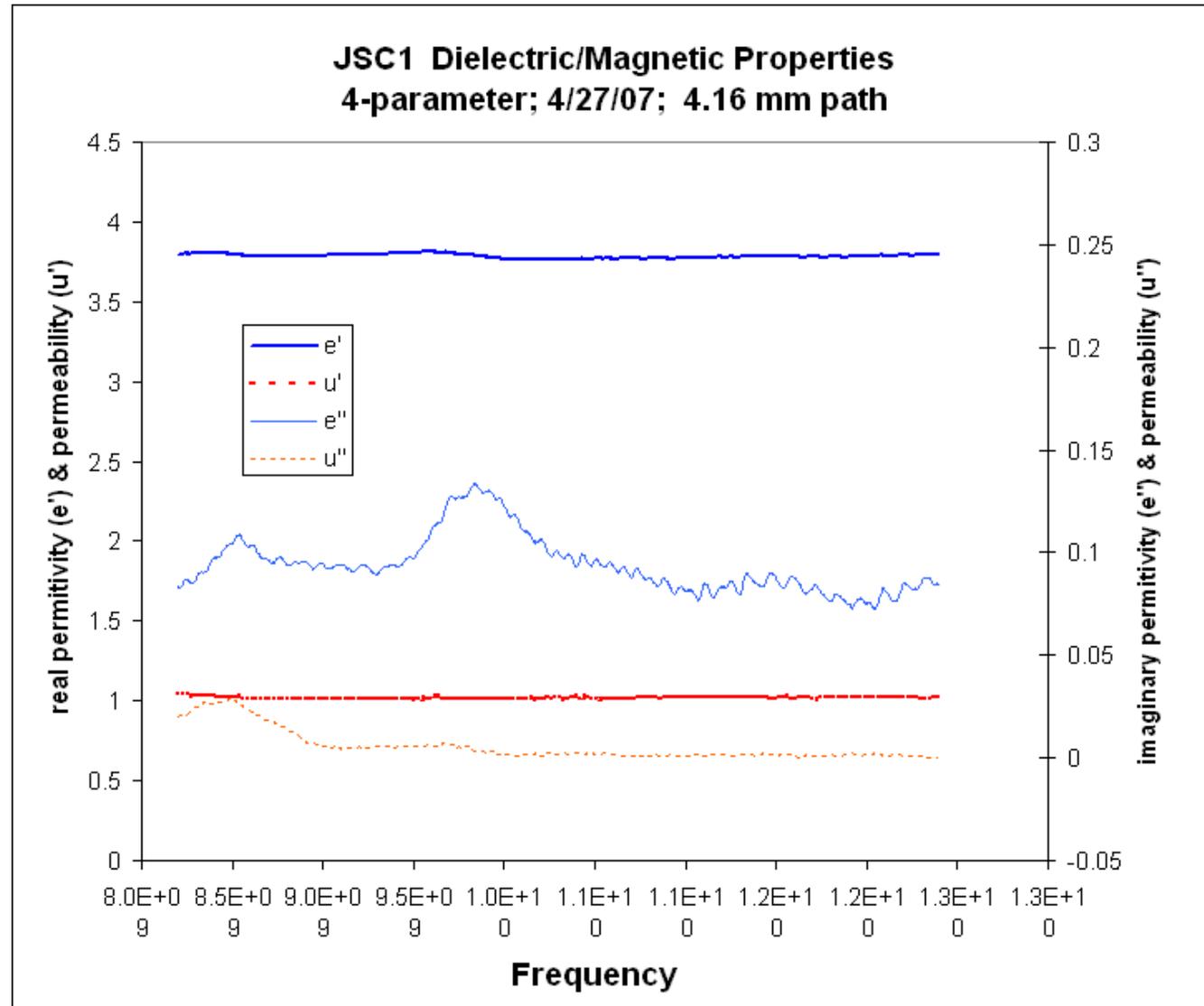
1. Electric Permittivity

(dielectric constant &  
Loss factor) and

2. Magnetic Permeability

We expect that Nano-  
phase Fe in lunar soil  
will significantly affect  
the permeability.

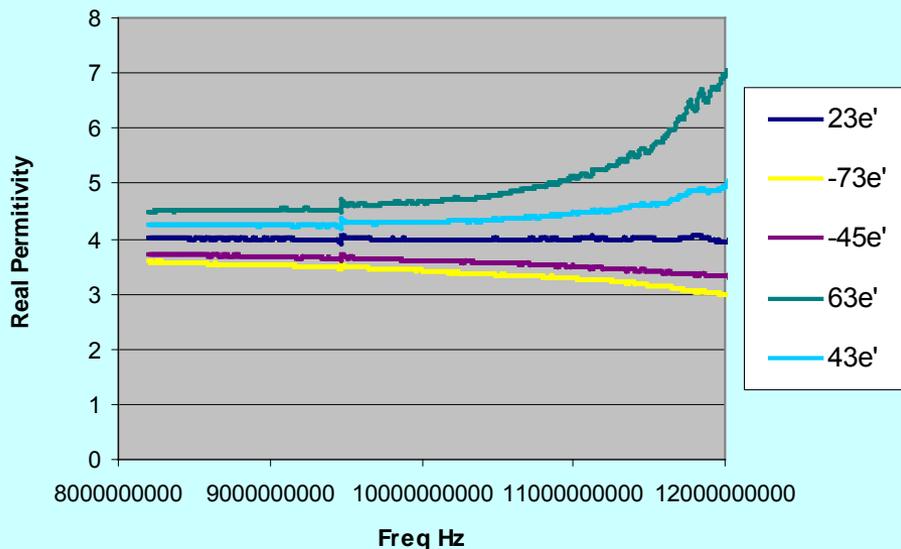
Proposal pending:  
Loan of **Apollo Soil**  
sample to measure  
dielectric properties.



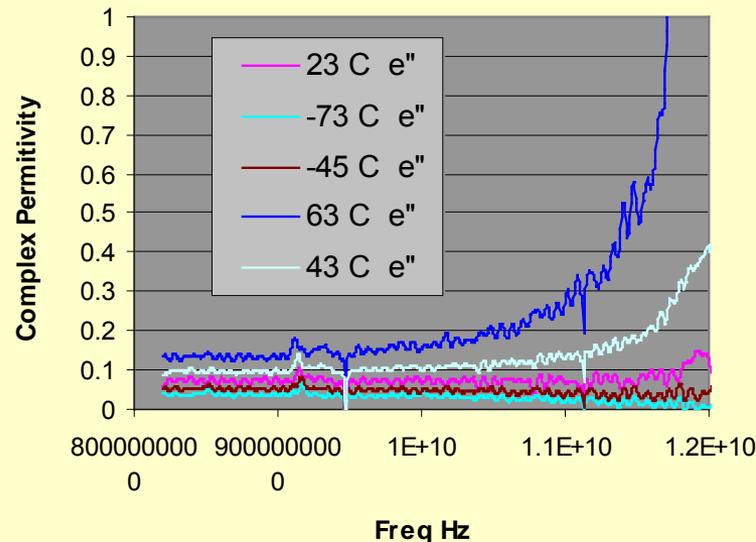
# Temperature Dependence (-73C to 63C) of the Dielectric Properties of Lunar Soil Simulant JSC-1A



JSC-1A Permittivity vs Temperature  
X-band, 8 to 12 GHz, Oct 30, 2007



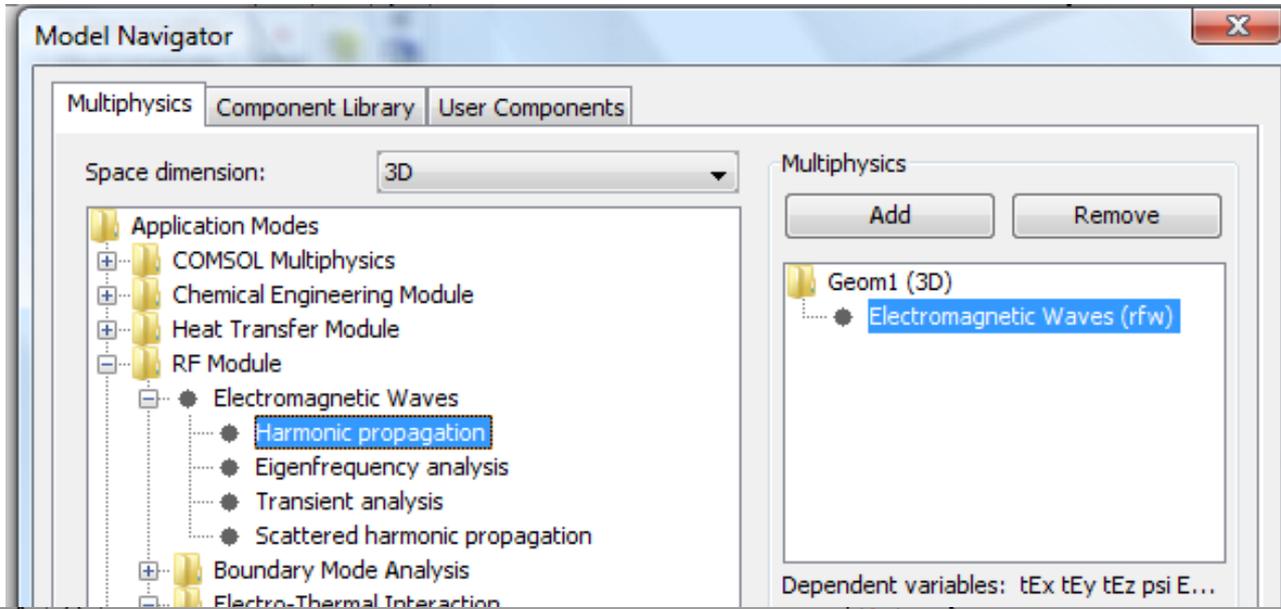
JSC-1A Complex Permittivity vs Temperature  
X-band, 8 to 12 GHz, Oct 30, 2007



Dielectric constant (real component of permittivity) vs. frequency over the X-band (8 to 12 GHz) for JSC-1A, temperatures from 63 C to -73 C.

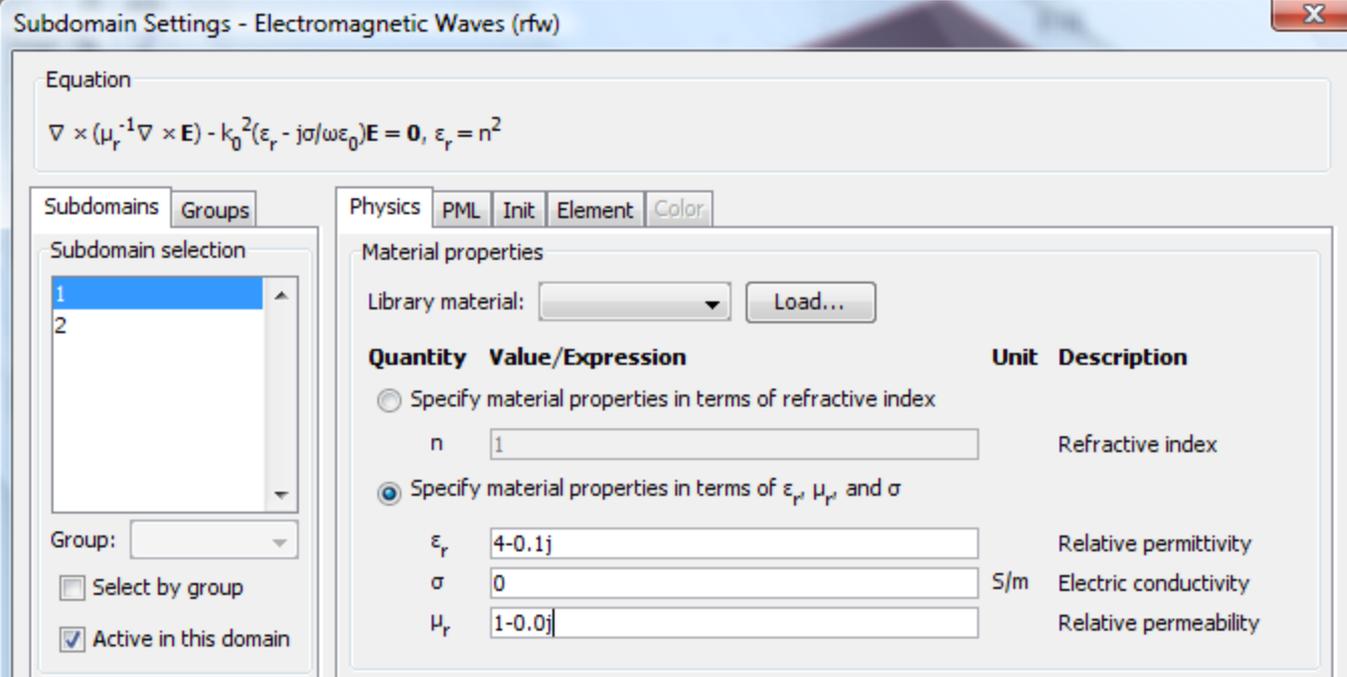
The loss factor (imaginary part of permittivity) vs. frequency 8 to 12 GHz 63 to -73 C.

# COMSOL - RF Module – Electromagnetic Waves – Harmonic Propagation



COMSOL 3.5a

Subdomain Settings



Electric Permittivity

4 - 0.1j

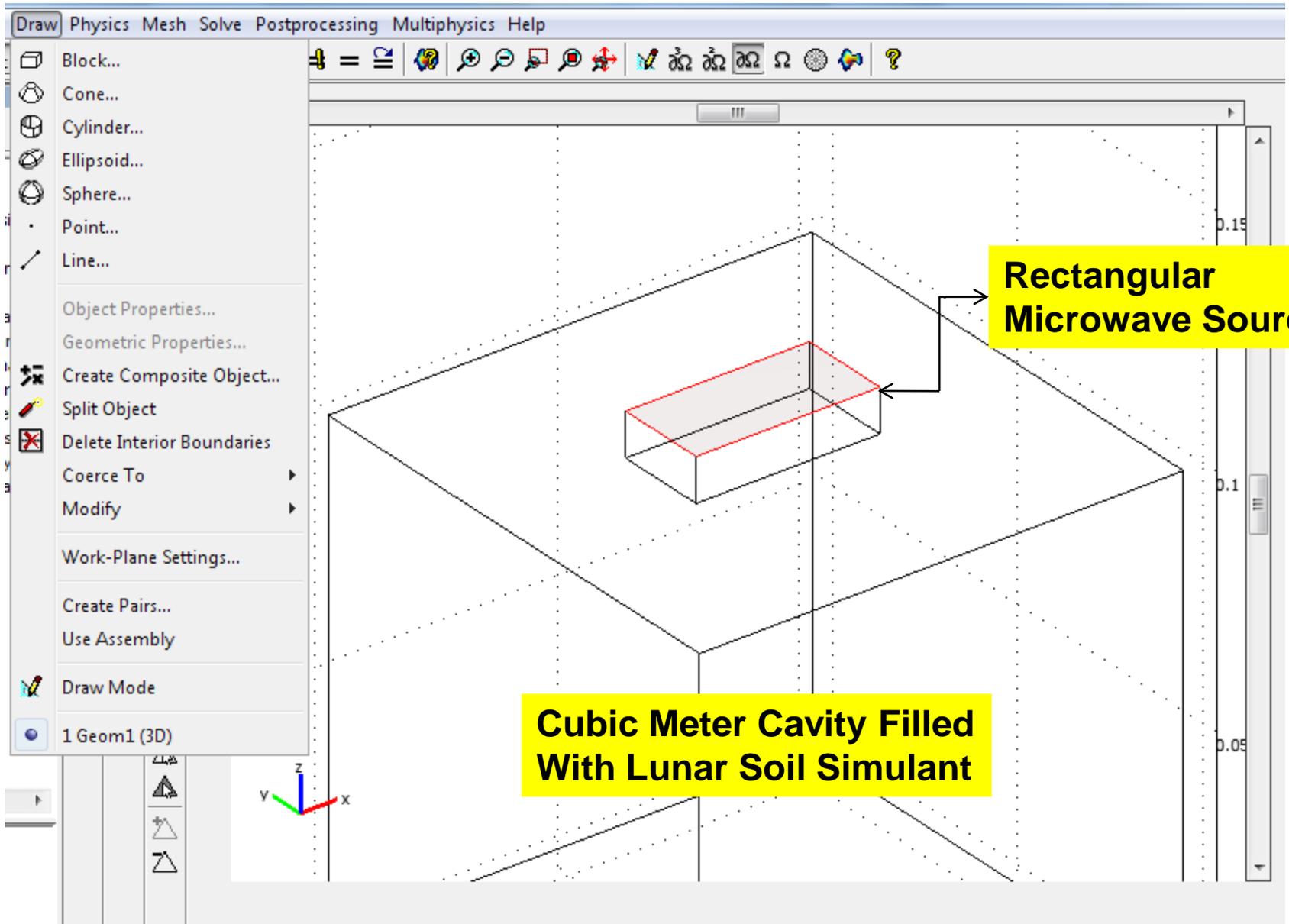
Magnetic Permeability

1 - 0.0j

Lunar Soil Simulant Dielectric Properties



# Geometry - Draw



# Boundary Conditions

Port

Rectangular Port

Continuity

Perfect Electric Conductor

Scattering Boundary

Conditions | Material Properties | Port | Far-Field | Color

Port definition

Mode specification: Rectangular

S-parameter output: Magnitude and phase

Mode type: Transverse electric (TE)

Mode number: 10

Port

Equation

$$S = \int (E - E_1) \cdot E_1 / \int E_1 \cdot E_1$$

Boundaries | Groups

Boundary selection

- 1 (volume)
- 2 (volume)
- 3 (volume)
- 4 (magnatron)
- 5 (volume)
- 6 (magnatron)
- 7 (magnatron)
- 8 (window)
- 9 (Port)

Conditions | Material Properties | Port | Far-Field | Color

Boundary sources and constraints

Boundary condition: Port

Port number: 1

Wave excitation at this port

Quantity	Value/Expression	Unit	Description
$P_{in}$	10	W	Port power level
$\Phi_p$	0		Port phase

Equation

$$\mathbf{n} \times (\nabla \times \mathbf{E}) - j\mathbf{k} \times (\mathbf{E} \times \mathbf{n}) = -\mathbf{n} \times (\mathbf{E}_0 \times j\mathbf{k}(\mathbf{n} - \mathbf{k})) \exp(-j\mathbf{k} \cdot \mathbf{r})$$

Boundaries | Groups

Boundary selection

- 1 (volume)
- 2 (volume)
- 3 (volume)
- 4 (magnatron)
- 5 (volume)
- 6 (magnatron)
- 7 (magnatron)
- 8 (window)
- 9 (Port)
- 10 (magnatron)

Conditions | Material Properties | Port | Far-Field | Color

Boundary sources and constraints

Boundary condition: Scattering boundary condition

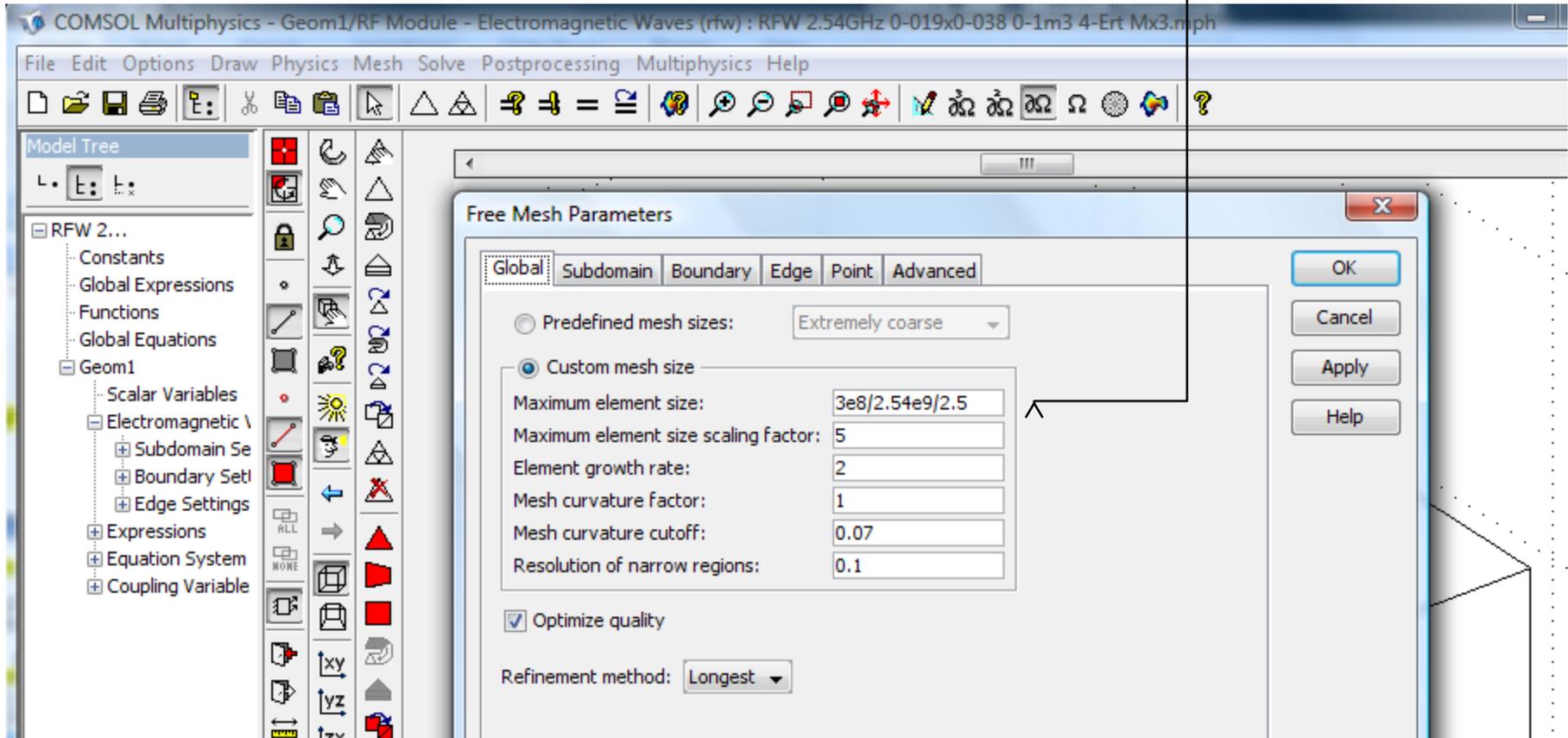
Quantity	Value/Expression	Unit	Description
Incident field:	Wave given by E field		
$E_0$	0 0 0	V/m	Electric field
Wave type:	Plane wave		
$\mathbf{k}$	-nx_rfw -ny_rfw -nz_rfw		Wave direction

Port Power  
10 Watt

# Meshing must satisfy the Nyquist criteria

$$\begin{aligned}\text{Maximum element size} &= c / \lambda / 2.5 \\ &= 3e8 / 2.45e9 / 2.5\end{aligned}$$

## Free Mesh Parameters

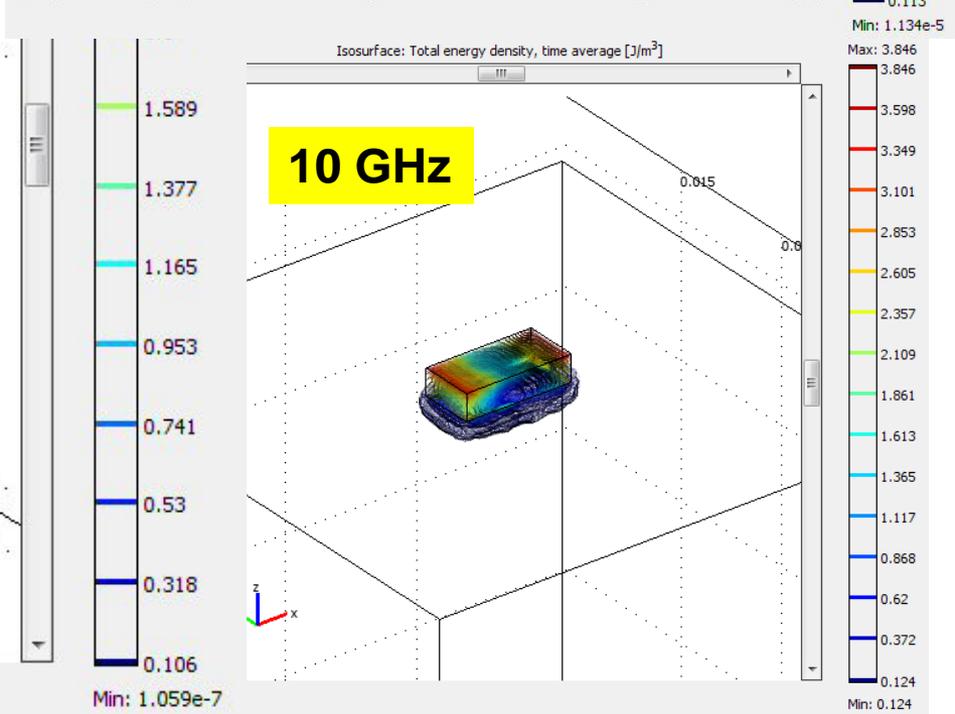
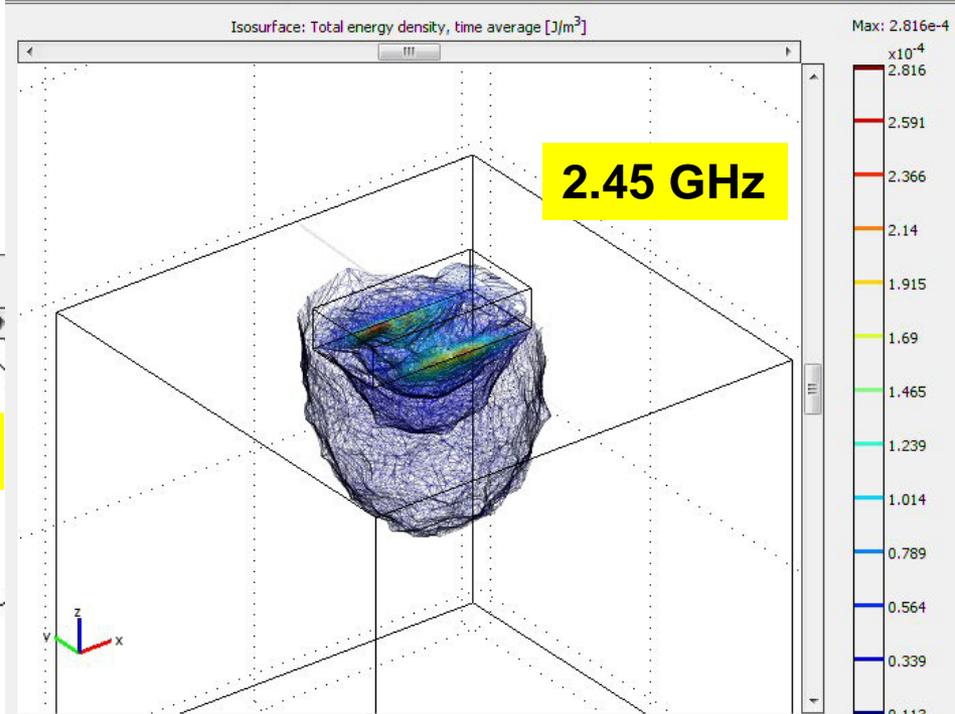
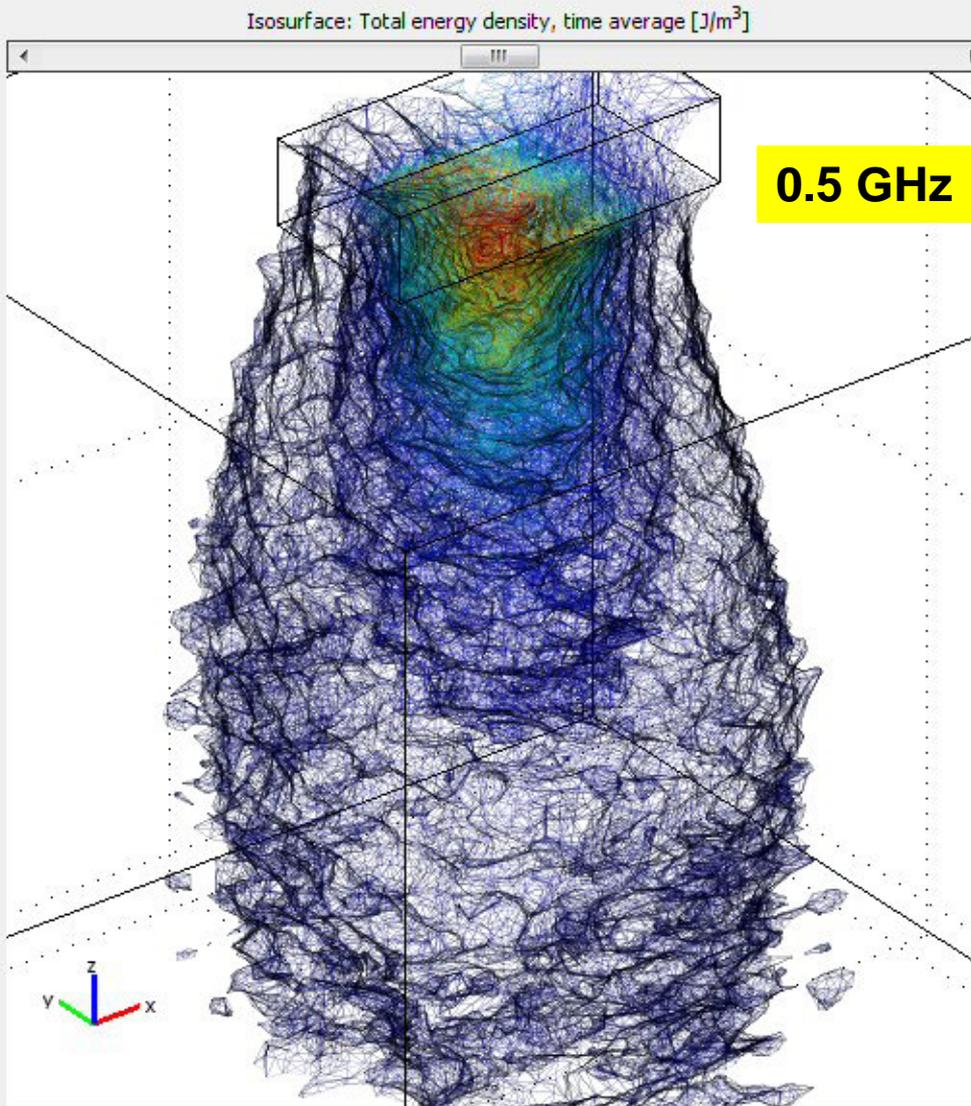


# Three Microwave Frequencies

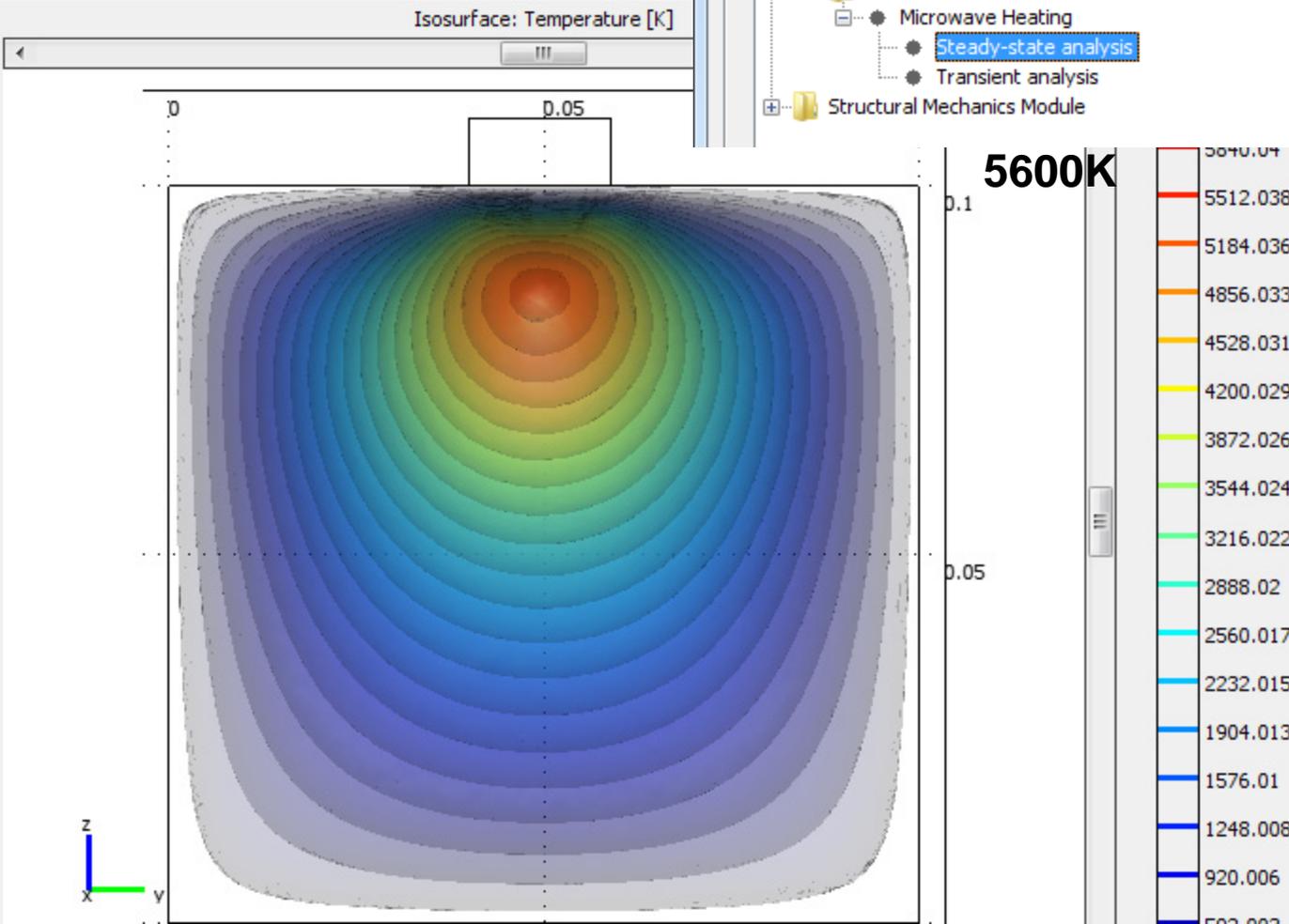
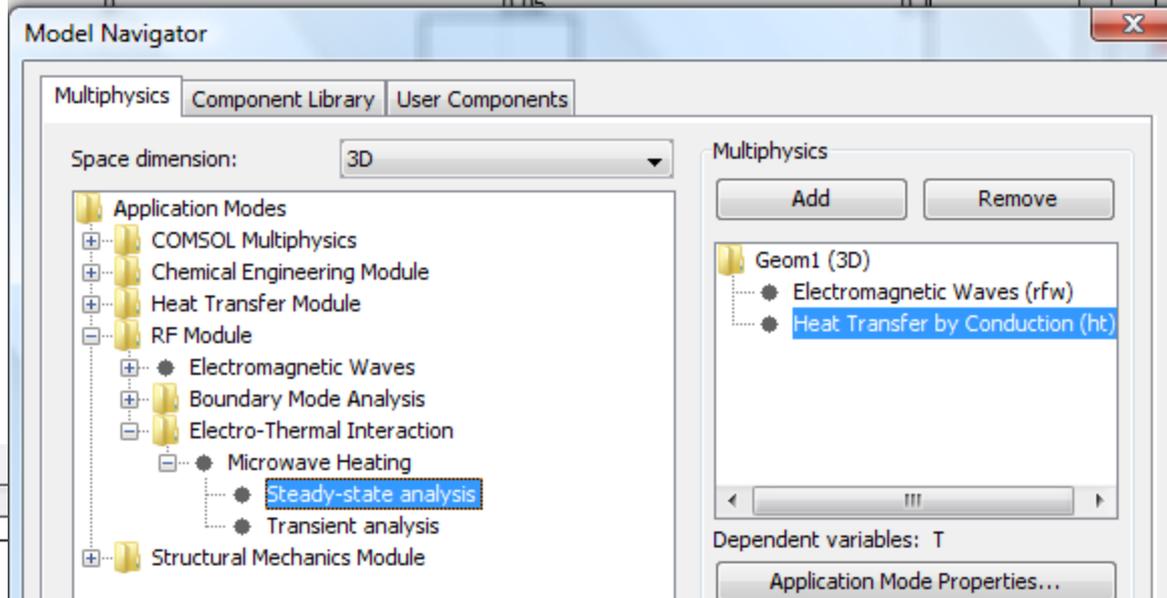


**Microwave flanges for the three different microwave frequencies (0.9 GHz, 2.45 GHz and 10 GHz) used in this project showing the relative sizes of the experimental and test measurement hardware. Their standard sizes are designated WR975, WR340 and WR90 respectively. Each microwave frequency requires different geometry for COMSOL.**

# RFW -Total Energy Density Iso-Surface Penetration into Simulant

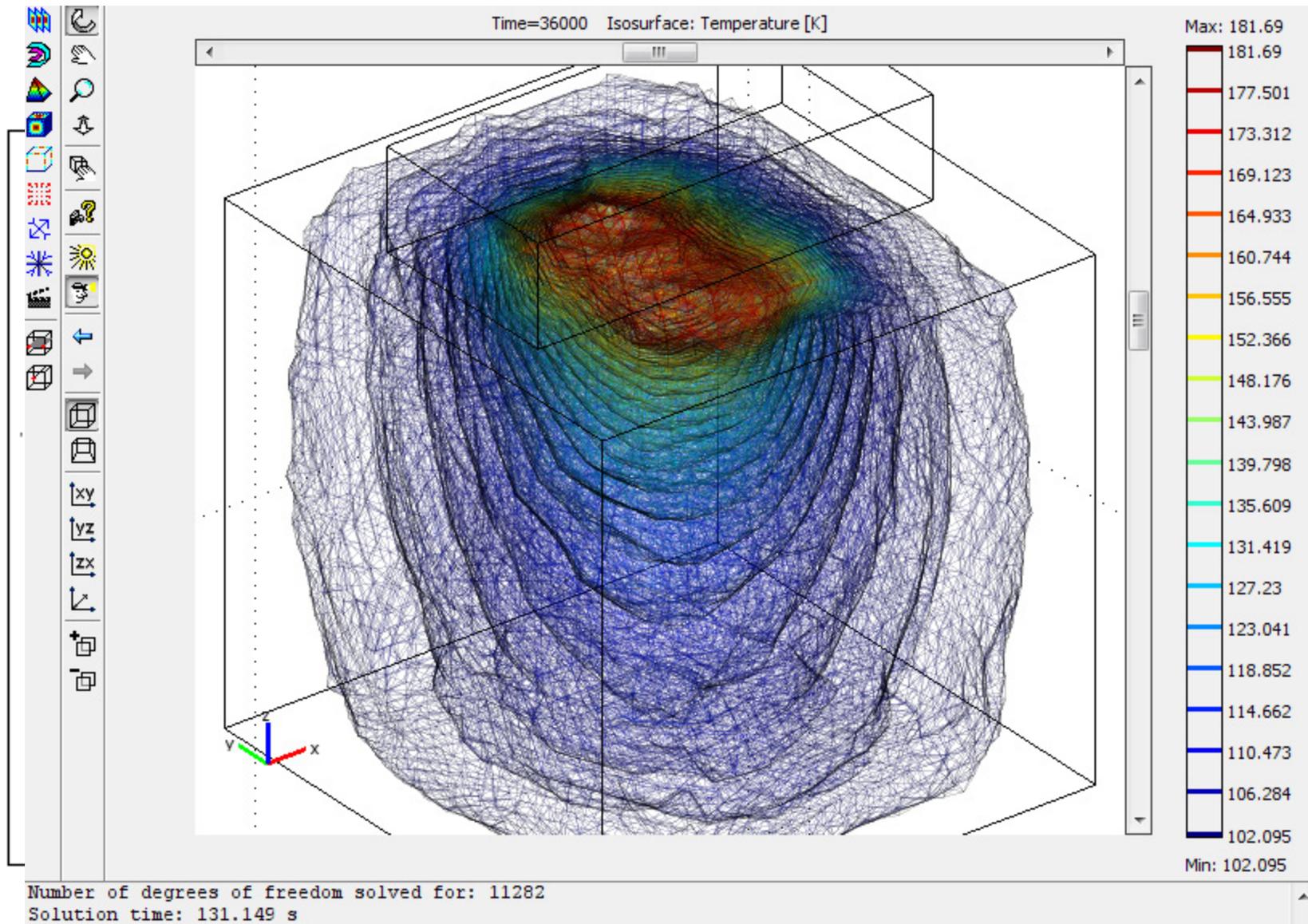


# Steady-State Electromagnetism Coupled with Heat Transfer by Conduction

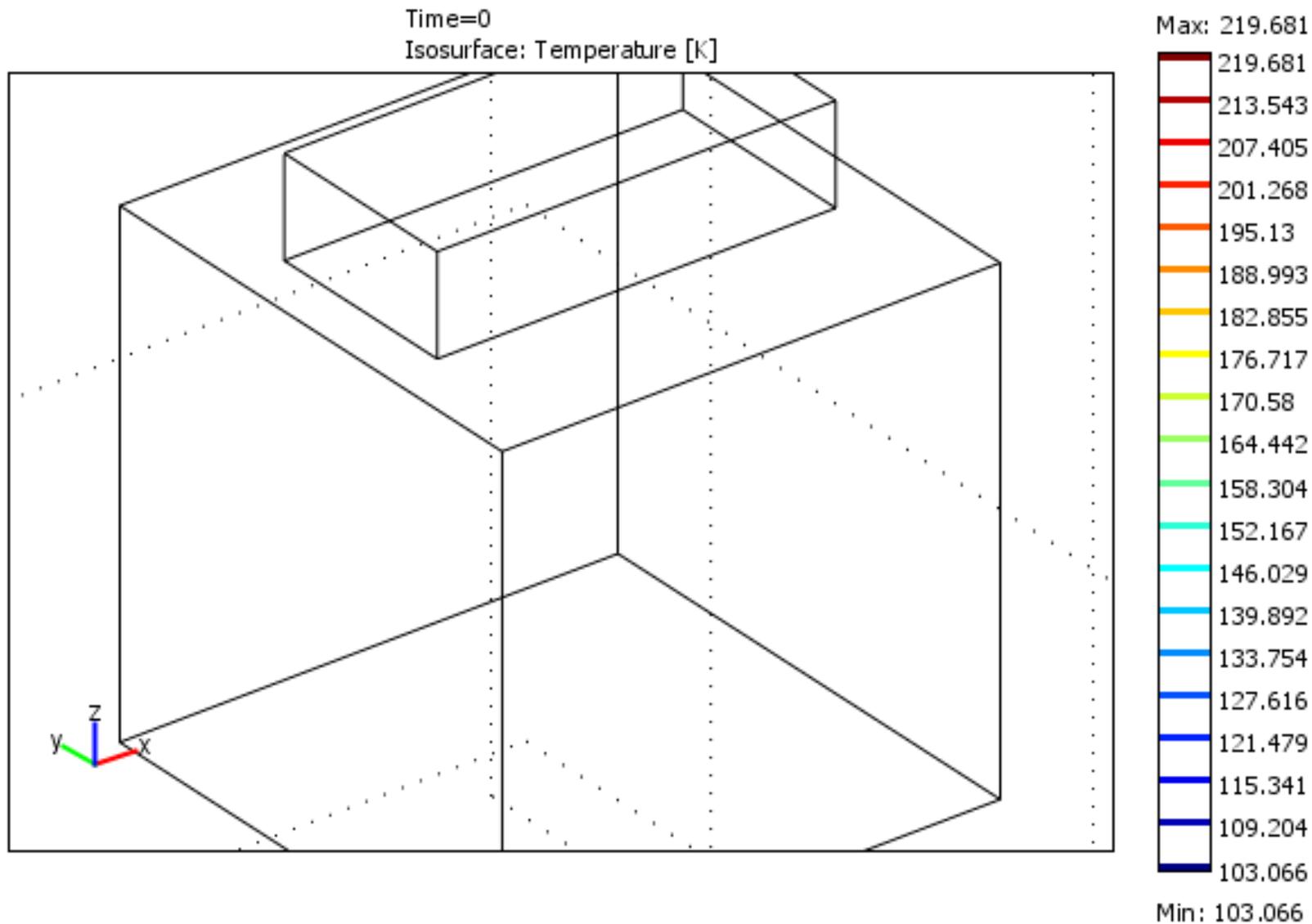


**2.45 GHz  
1 KW power  
Temperatures to  
5000K**

# RFW coupled with Heat Conduction - Transient Temperature Isotherms, 10 hours Click to start animation



# RFW coupled with Heat Conduction - Transient Temperature Isotherms, 10 hours



# Application of COMSOL

- **Processing parameters and hardware requirements for water extraction is a complex multi-physics problem.**
- **Microwave coupling to materials and heating is dependent on frequency and materials properties.**
- **Materials properties are a function of frequency and temperature.**
- **Can calculate microwave penetration and heating, with frequency and temperature dependent lunar soil dielectric properties.**
- **To Do – Model the Percolation of water vapor through the soil (porous media) characterized by the Darcy constant (currently being measured by Southern Research Institute).**
- **Parametric modeling will permit the evaluation of processing parameters most suitable for prototype hardware development, testing, and trade studies.**

# Acknowledgements:

**NASA HQ** – Joint Science Mission Directorate (**SMD**) & Exploration Technology Development Program (**ETDP**) Lunar Advanced Science and Exploration Research (**LASER**) program

**MSFC Management – Prior Seed Funding**

**Co-Investigators:**

**Dr. William Kaukler** – University of Alabama-Huntsville  
Hardware Instrumentation  
Dielectric Property Test Configurations

**Frank Hepburn** – Materials & Processes, MSFC-NASA  
Dielectric Property Measurements - Network Analyzer

**COMSOL – Walter Frei** -Technical assistance with the models

## Contacts:

**MSFC Public Affairs** – [Steve.Roy@msfc.nasa.gov](mailto:Steve.Roy@msfc.nasa.gov)

**MSFC Technology Licensing** – 256-544-5353

[ed.ethridge@nasa.gov](mailto:ed.ethridge@nasa.gov)

# END